

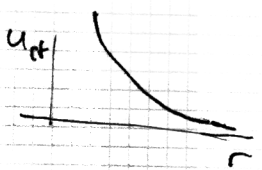
8.14 *** Consider a particle of reduced mass μ orbiting in a central force with $U = kr^n$ where $kn > 0$.
 (a) Explain what the condition $kn > 0$ tells us about the force. Sketch the effective potential energy U_{eff} for the cases that $n = 2, -1$, and -3 . (b) Find the radius at which the particle (with given angular momentum ℓ) can orbit at a fixed radius. For what values of n is this circular orbit stable? Do your sketches confirm this conclusion? (c) For the stable case, show that the period of small oscillations about the circular orbit is $\tau_{\text{osc}} = \tau_{\text{orb}}/\sqrt{n+2}$. Argue that if $\sqrt{n+2}$ is a rational number, these orbits are closed. Sketch them for the cases that $n = 2, -1$, and 7 .

$U = kr^n, kn > 0$

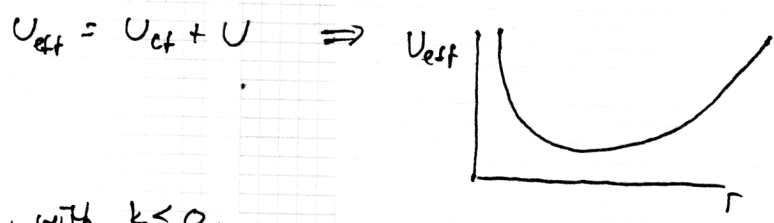
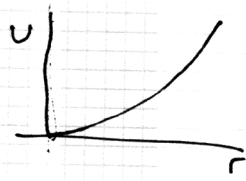
(a) What does kn indicate about force?

$F = -\frac{\partial U}{\partial r} = -knr^{n-1} \Rightarrow kn > 0 \Rightarrow F(r) < 0$, i.e. force is in $-\hat{r}$ direction, so is attractive.

$U_{\text{cf}} = \frac{\ell^2}{2\mu r^2}$

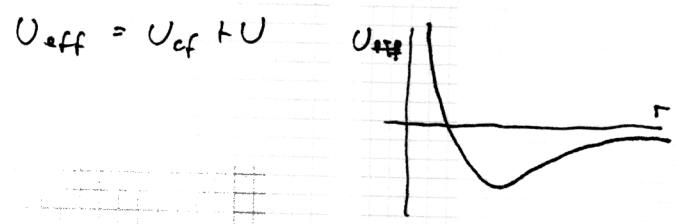
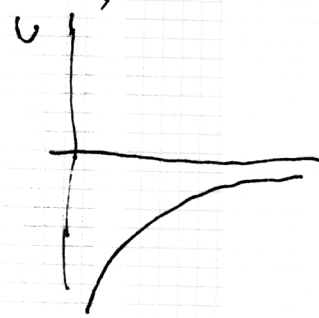


(i) for $n = 2, U = kr^2$

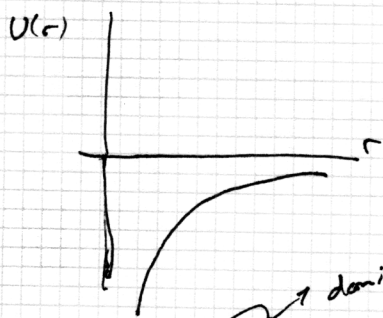


(ii) for $-1 = n$, with $k < 0$,

$U = -\frac{|k|}{r}$

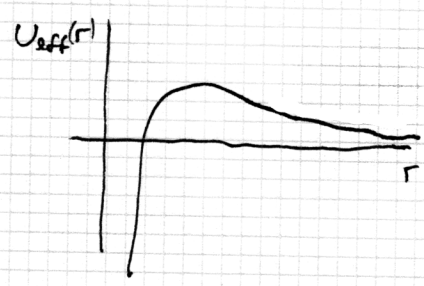


(iii) $n = -3, k < 0$



$U_{\text{eff}} = -\frac{k}{r^3} + \frac{l_z^2}{2\mu r^2}$

Annotations:
 - Arrow pointing to the $-\frac{k}{r^3}$ term: "dominates for smaller r"
 - Arrow pointing to the $\frac{l_z^2}{2\mu r^2}$ term: "dominates for larger r"



(b) $U_{\text{eff}} = U(r) + \frac{l_z^2}{2\mu r^2} = kr^n + \frac{l_z^2}{2\mu r^2}$

radial eqn: $m\ddot{r} = -\frac{\partial U_{\text{eff}}}{\partial r}$

→ want values of r for which $r = \text{const.}$ For these values of r, will have $\ddot{r} = 0$, so want solns to

$$\frac{\partial U_{\text{eff}}}{\partial r} \Big|_{r=r_0} = 0$$

$$\frac{\partial U_{\text{eff}}}{\partial r} \Big|_{r=r_0} = knr_0^{n-1} - \frac{l_z^2}{\mu r_0^3} = 0$$

$$knr_0^{n-1} = \frac{l_z^2}{\mu r_0^3}$$

$$r_0^{n+2} = \frac{l_z^2}{kn\mu}$$

$$r_0 = \left[\frac{l_z^2}{kn\mu} \right]^{\frac{1}{n+2}}$$

(assume $kn > 0$)

→ For what values of n is this stable?

Orbit is stable if $\frac{\partial^2 U}{\partial r^2} \Big|_{r=r_0} > 0.$

$$\frac{\partial^2 U_{\text{eff}}}{\partial r^2} = \frac{\partial}{\partial r} \left[knr^{n-1} - \frac{l_z^2}{\mu r^3} \right] = kn(n-1)r^{n-2} + \frac{3l_z^2}{\mu r^4}$$

at $r=r_0$,

$$\left. \frac{d^2 U_{\text{eff}}}{dr^2} \right|_{r=r_0} = kn(n-1)r_0^{n-2} + \frac{3l_z^2}{\mu r_0^4} > 0$$

multiply by r_0^4 :

$$kn(n-1)r_0^{n+2} + \frac{3l_z^2}{\mu} > 0$$

$$\cancel{kn(n-1)} \frac{l_z^2}{k\mu} + \frac{3l_z^2}{\mu} > 0$$

$$(n-1) + 3 > 0$$

$$n+2 > 0$$

 $n > -2$ for stable orbits.

This is consistent w/ sketches:

 $n=2, -1$ have minima of U at $r=r_0$ for $n=-3$ it's a local max at $r=r_0$.

© Show that period of small oscillations about circular orbit is:

$$\tau_{\text{osc}} = \frac{\tau_{\text{orb}}}{\sqrt{n+2}}$$

radial equation: $\mu \ddot{r} = -\frac{\partial U_{\text{eff}}(r)}{\partial r}$ For small perturbations about r_0

$$r = r_0 + \epsilon$$

$$\mu \ddot{r} \rightarrow \mu \ddot{\epsilon}$$

$$\frac{\partial U_{\text{eff}}}{\partial r} \approx \frac{\partial U_{\text{eff}}(r_0)}{\partial r} + \epsilon \frac{\partial^2 U_{\text{eff}}(r_0)}{\partial r^2}$$

$$\text{eqn of motion is: } \mu \ddot{\epsilon} + \epsilon \frac{\partial^2 U_{\text{eff}}(r_0)}{\partial r^2} = 0$$

$$\begin{aligned} \frac{\partial^2 U_{\text{eff}}(r_0)}{\partial r^2} &= kn(n-1)r_0^{n-2} + \frac{3l_z^2}{\mu r_0^4} = \frac{1}{r_0^4} \left(kn(n-1)r_0^{n+2} + \frac{3l_z^2}{\mu} \right) \\ &= \frac{1}{r_0^4} \left[k\mu(n-1) \frac{l_z^2}{k\mu} + \frac{3l_z^2}{\mu} \right] = \frac{l_z^2}{\mu} \frac{1}{r_0^4} (n+2) \end{aligned}$$

$$\ddot{\epsilon} + \frac{\epsilon}{\mu} \frac{\partial^2 U_{\text{eff}}(r_0)}{\partial r^2} = 0$$

$$\ddot{\epsilon} + \omega_0^2 \epsilon = 0, \quad \omega_0^2 = \frac{1}{\mu} \frac{\partial^2 U_{\text{eff}}(r_0)}{\partial r^2} = \frac{l_z^2}{\mu^2} \frac{n+2}{r_0^4}$$

$$\text{and } \tau_{\text{rel}} = \frac{2\pi}{\omega_0} = 2\pi \frac{\mu}{l_z} \frac{r_0^2}{\sqrt{n+2}}$$

Period of angular motion?

8-14-4

$$\dot{\phi} = \frac{l_z}{\mu r_0^2} \Rightarrow \frac{\Delta\phi}{\Delta t} = \frac{l_z}{\mu r_0^2} \cdot \text{For } \Delta\phi = 2\pi, \text{ we get } \Delta t = \tau.$$

using $r_0 = \left[\frac{l_z^2}{k\mu} \right]^{\frac{1}{n+2}}$,

so: $\tau_{\text{ang}} = \frac{2\pi \mu r_0^2}{l_z}$

we get $\tau_{\text{ang}} = \frac{2\pi \mu r_0^2}{l_z}$

so: $\tau_{\text{rad}} = \frac{\tau_{\text{ang}}}{\sqrt{n+2}}$ (i.e. $\tau_{\text{osc}} = \frac{\tau_{\text{orb}}}{\sqrt{n+2}}$)

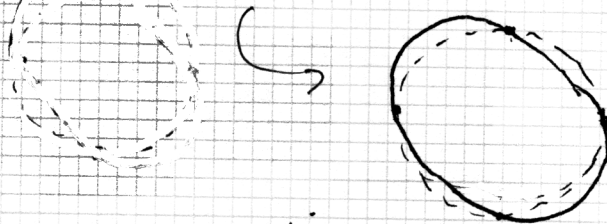
- If $\sqrt{n+2}$ is rational, call it $\frac{p}{q}$, we have

$\tau_{\text{rad}} = \frac{q}{p} \tau_{\text{ang}}$. After a time $p\tau_{\text{rad}} = q\tau_{\text{ang}} = T$, the angular position and the radial position are the same as the initial positions and the orbit closes.

- $n=2$: $\tau_{\text{osc}} = \frac{\tau_{\text{orb}}}{2} \Rightarrow \tau_{\text{orb}} = 2\tau_{\text{osc}}$
- $n=-1$: $\tau_{\text{osc}} = \tau_{\text{orb}}$
- $n=7$: $\tau_{\text{osc}} = \frac{\tau_{\text{orb}}}{3} \Rightarrow \tau_{\text{orb}} = 3\tau_{\text{osc}}$

$n=2$: 2 oscillations per orbit

(so oscillating path crosses circular orbit 4 times per orbit)



$n=4$: 1 osc. per orbit (osc. path crosses circular orbit twice per orbit)



$n=7$: 3 osc. per orbit (osc. path crosses circ. orbit 6 times per orbit)

